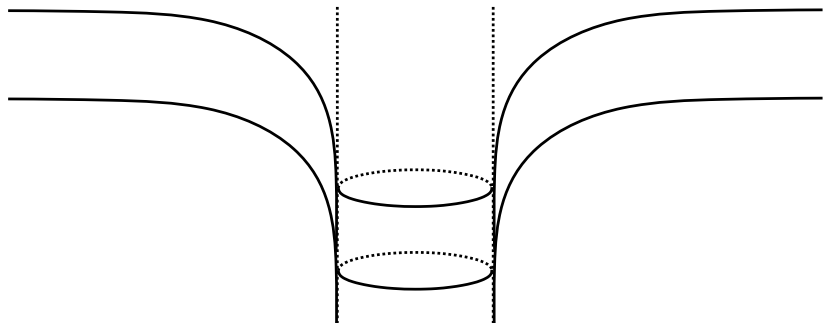


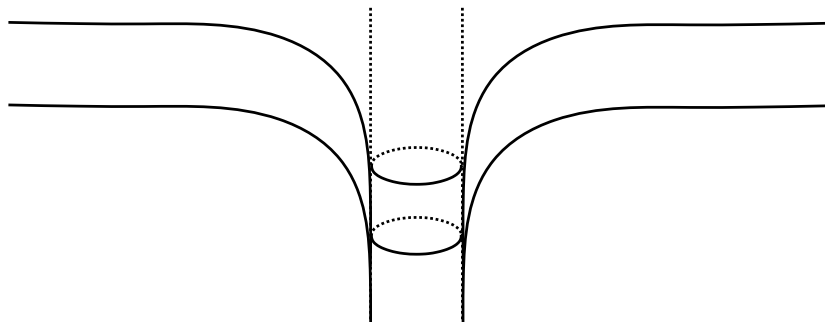
Uniform Temple charts and applications to null distance

Benjamin Meco,
joint with Anna Sakovich and Christina Sormani

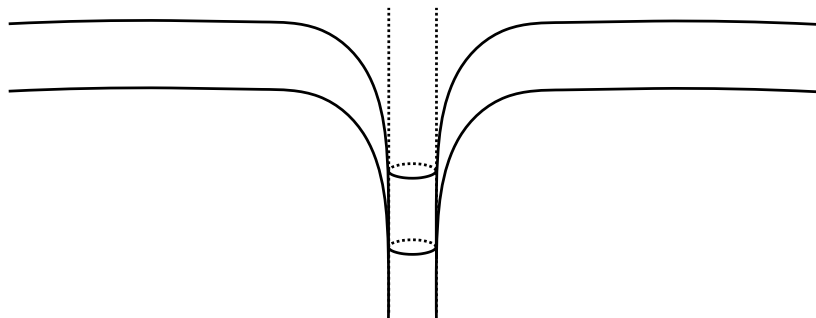
Convergence of spacetimes



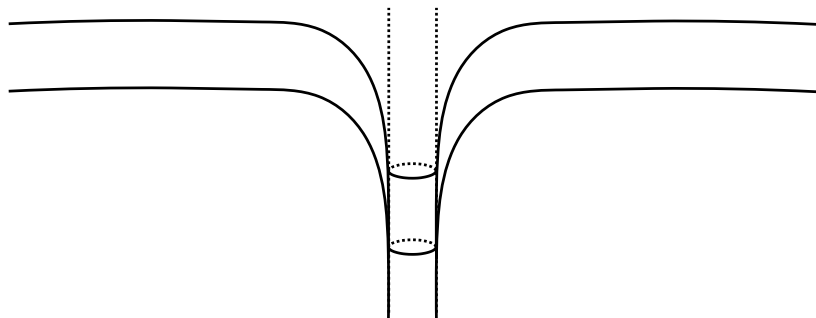
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Convergence of spacetimes



We would like a notion of distance between spacetimes that applies when spacetimes are not diffeomorphic and their metric tensors are not close in a smooth sense.

Why do we need a metric theory of spacetimes?

Other questions one could try to answer:

- ▶ Is our spacetime well-approximated by cosmological models?
- ▶ Can we compare spacetimes with different topology?
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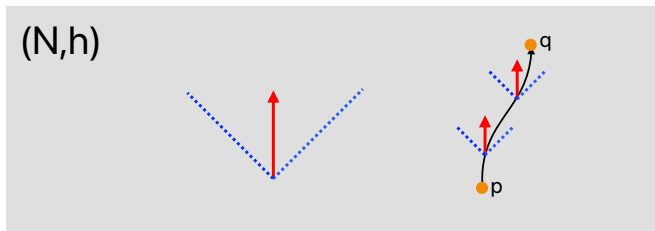
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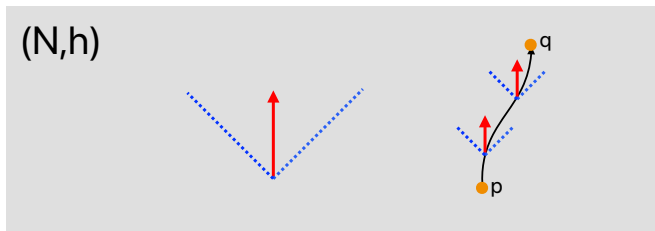
Problem: In contrast to Riemannian manifolds spacetimes are not natural metric spaces.

Spacetimes and causal structure



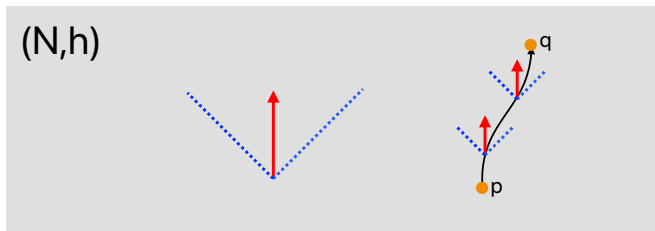
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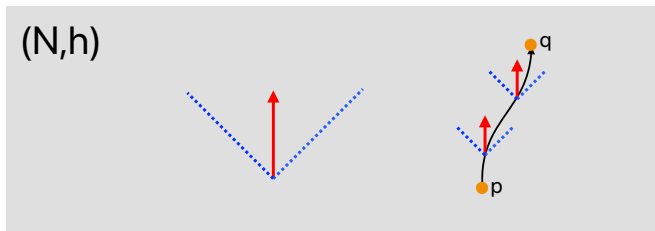
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- ▶ A curve $\gamma : I \rightarrow N$ is causal if $\dot{\gamma}$ is always causal.
- ▶ We denote the causal future of a point $p \in N$ by $J^+(p)$.

Time functions

Definition (Time functions)

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Could be infinite, a **regular** cosmological time function takes values in $(0, \infty)$.

The null distance of Sormani and Vega

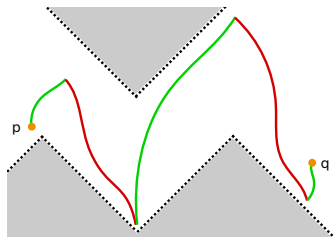
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where γ is a piecewise causal curve from p to q , with breakpoints at each s_1, \dots, s_{k-1} .

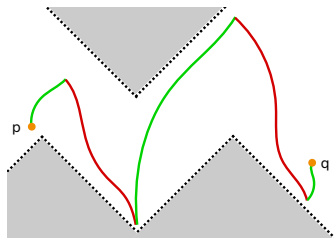


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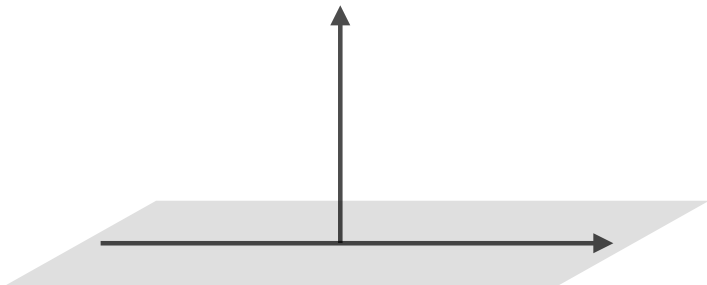
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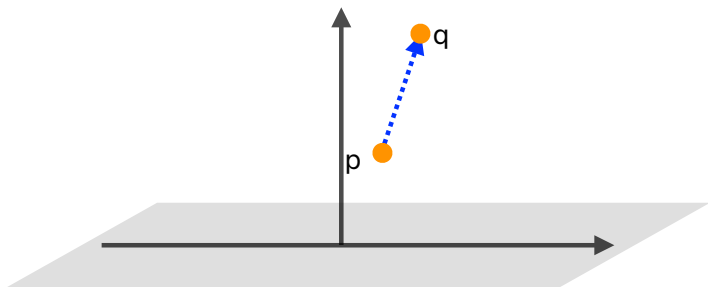


If τ is **locally anti-Lipschitz** (Chruściel-Grant-Minguzzi 2016), e.g. $\tau = \tau_g$ is a regular cosmological time function, then \hat{d}_τ is definite.

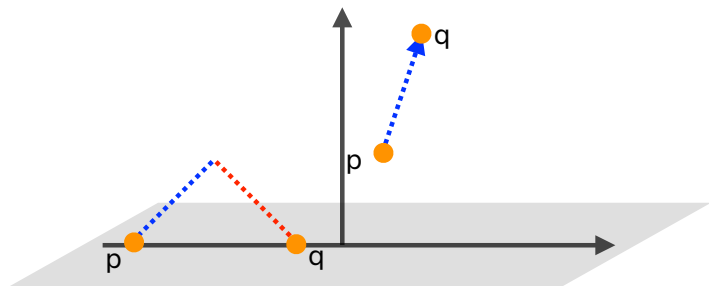
Example: Null distance in Minkowski space, $\tau = t$



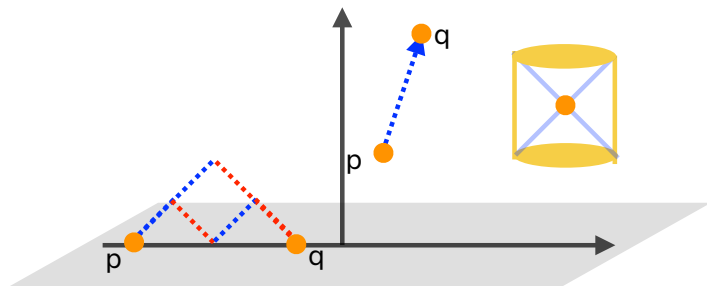
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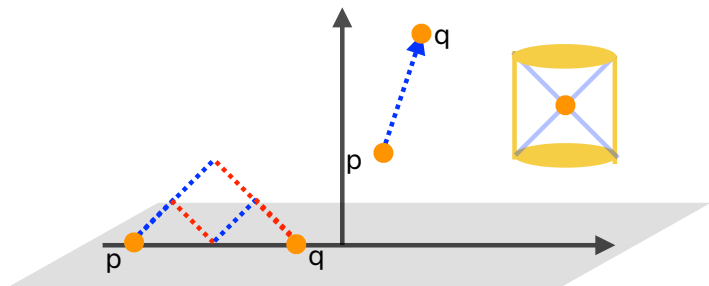
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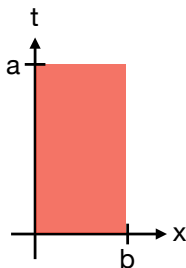
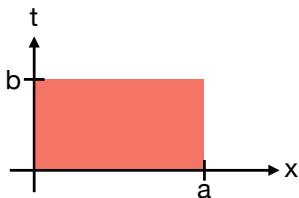
i.e. spacetimes are "the same" iff there is a Lorentzian isometry between them.

Example: A distance preserving map which does not preserve causal structure

Consider the following subsets of $(\mathbb{R}^2, -dt^2 + dx^2)$, $\tau(t, x) = t$.

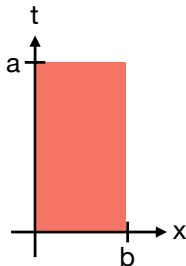
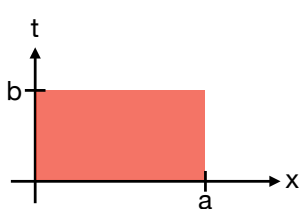
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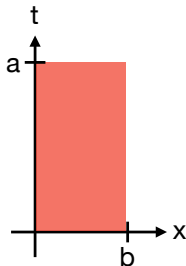
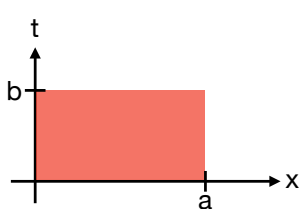


In both cases, the induced null distance is

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The map $F(t, x) = (x, t)$ is a metric isometry, but these spacetimes are not isometric as Lorentzian manifolds.

Encoding causality

We say that τ and \hat{d}_τ **encode causality** if for all $p, q \in N$:

$$\hat{d}_\tau(p, q) = \tau(q) - \tau(p) \iff q \in J^+(p).$$

This has been proved to hold:

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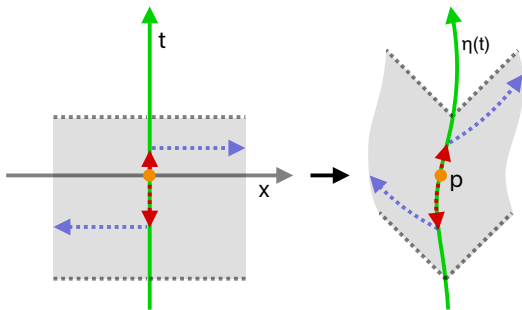
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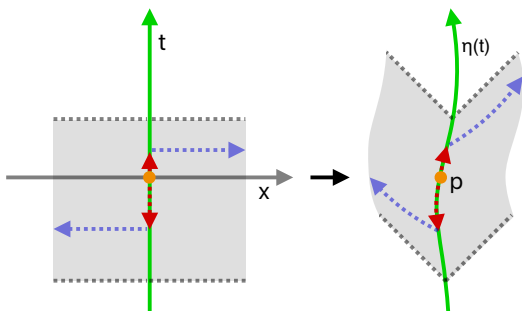
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Temple charts



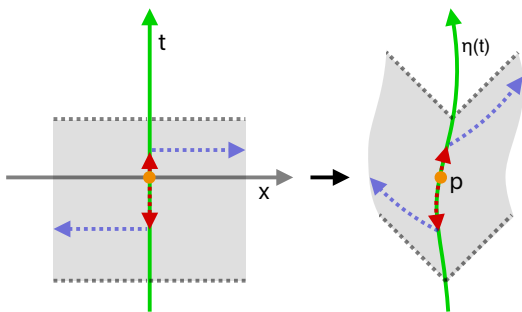
Temple charts



$$\Phi_p(t, x^1, \dots, x^n) = \exp_{\eta(t)} \left(|\vec{x}| e_0 + \sum_{i=1}^n x^i e_i \right), \quad \omega_p(\Phi_p(t, \vec{x})) := t,$$

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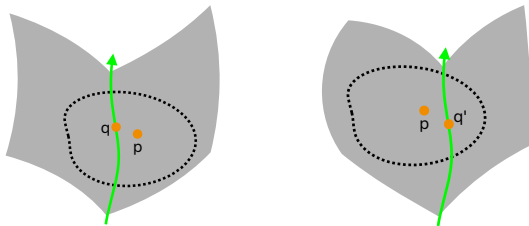
where $\{e_0, \dots, e_n\}$ are frame fields and η is a future timelike geodesic with $\eta(0) = p$ and $\dot{\eta} = e_0$. We have:

$$q \in J^+(p) \iff \omega_p(q) \geq 0.$$

Uniform Temple charts

Theorem (M.-Sakovich-Sormani, TBP)

Every $p \in N$ has a neighborhood U_p such that for all $q \in U_p$, the image of the Temple chart Φ_q covers U_p .

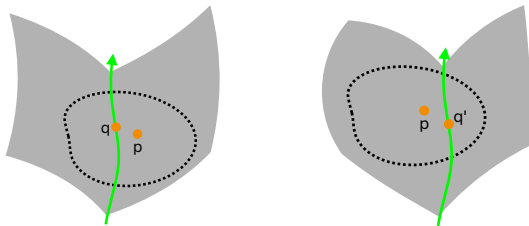


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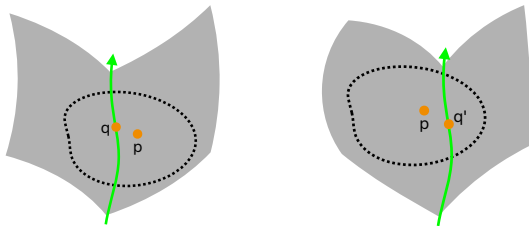


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Moreover, there is a Riemannian metric, g_R , on U_p such that

$$\text{for all } q, q' \in U_p: K^{-1}d_{g_R}(q, q') \leq \hat{d}_\tau(q, q') \leq Kd_{g_R}(q, q').$$

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Remark: This was proven in Sakovich-Sormani 2022 under an extra assumption that causality is globally encoded by \hat{d}_{τ_i} and τ_i .

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- ▶ the time function τ and the null distance \hat{d}_τ locally encode causality,
- ▶ the "converted" metric space (N, \hat{d}_τ) can be equipped with a bi-Lipschitz atlas,
- ▶ a time function and null distance preserving bijection must be a smooth Lorentzian isometry.

Sakovich and Sormani 2024 discuss several possible definitions of (definite) distances between spacetimes, based on τ and \hat{d}_τ .

Thank you for listening!