

# Developments of initial data on big bang singularities for the Einstein–nonlinear scalar field equations

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Andrés Franco Grisales January 23, 2025 — 15th Central European Relativity Seminar



Fournodavlos and Luk (2020) constructed solutions to the Einstein vacuum equations of the form

$$g = -dt \otimes dt + \sum_{i,j=1}^{3} a_{ij} t^{2p_{\max\{i,j\}}} dx^{i} \otimes dx^{j}$$

on  $(0, T] \times \mathbb{T}^3$ , where  $p_i : \mathbb{T}^3 \to \mathbb{R}$  and  $a_{ij} : (0, T] \times \mathbb{T}^3 \to \mathbb{R}$  are smooth and satisfy

 $\lim_{t\to 0^+}a_{ij}(t)=c_{ij},$ 

where  $c_{ij} : \mathbb{T}^3 \to \mathbb{R}$  are some prescribed smooth functions.



Constructions of quiescent big bang spacetimes:

- In symmetric settings: Isenberg, Moncrief, 1990, 2002; Kichenassamy, Rendall, 1998; Rendall, 2000; Ståhl, 2002; Choquet-Bruhat, Isenberg, Moncrief, 2004; Ames, Beyer, Isenberg and LeFloch, 2013, 2017.
- Analytic spacetimes without symmetry: Andersson, Rendall, 2001; Damour, Henneaux, Rendall, Weaver, 2002; Klinger, 2015.
- No symmetry or analyticity: Fournodavlos, Luk, 2020; Athanasiou, Fournodavlos, 2024.

Ringström (2022) introduced a geometric notion of initial data on big bang singularities.



Let (M, g) be a 4-dimensional spacetime,  $\varphi \in C^{\infty}(M)$  and  $V \in C^{\infty}(\mathbb{R})$ . The equations are

$${
m Ric} - rac{1}{2} Sg + \Lambda g = T$$
 , $\Box_g arphi = V' \circ arphi$ 

where T is given by

$$T = d\varphi \otimes d\varphi - \left(rac{1}{2}|d\varphi|_g^2 + V \circ \varphi
ight)g.$$



Initial data on the singularity

## Definition (Ringström '22)

Let  $(\Sigma, \mathring{\mathcal{H}})$  be a closed 3-dimensional Riemannian manifold,  $\mathring{\mathcal{K}}$  a (1,1)-tensor on  $\Sigma$  and  $\mathring{\Phi}, \mathring{\Psi} \in C^{\infty}(\Sigma)$  such that:

1. tr $\mathring{\mathcal{K}} = 1$  and  $\mathring{\mathcal{K}}$  is symmetric with respect to  $\mathring{\mathcal{H}}$ .



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4.  $\gamma_{23}^1 = 0$  in a neighborhood of  $x \in \Sigma$  if  $p_1(x) \leq 0$ , where  $[e_i, e_k] = \gamma_{ik}^{\ell} e_{\ell}$  and  $\mathring{\mathcal{K}}(e_i) = p_i e_i$  with  $p_1 < p_2 < p_3$ .

Then  $(\Sigma, \mathring{\mathcal{H}}, \mathring{\mathcal{K}}, \mathring{\Phi}, \mathring{\Psi})$  are initial data on the singularity.



## The expansion normalized quantities

#### Definition

Let (M, g) be a spacetime and  $\Sigma \subset M$  a spacelike hypersurface with future pointing unit normal U. Let h and k denote the induced metric and the second fundamental form of  $\Sigma$ . If tr<sub>h</sub>k > 0,

$$\mathcal{K}(X) := \frac{1}{\mathrm{tr}_h k} k(X, \cdot)^{\sharp}, \qquad \mathcal{H}(X, Y) := h\Big((\mathrm{tr}_h k)^{\mathcal{K}}(X), (\mathrm{tr}_h k)^{\mathcal{K}}(Y)\Big),$$
  
where  $(\mathrm{tr}_h k)^{\mathcal{K}}(X) := \sum_{n=0}^{\infty} \frac{(\ln \mathrm{tr}_h k)^n}{n!} \mathcal{K}^n(X)$  and  $X, Y \in \mathfrak{X}(\Sigma).$   
For  $\varphi \in C^{\infty}(M),$   
 $\Psi := \frac{1}{\mathrm{tr}_h k} U \varphi, \qquad \Phi := \varphi + \Psi \ln(\mathrm{tr}_h k).$ 



## Development of the data

### Definition

Let  $(\Sigma, \mathcal{H}, \mathcal{K}, \Phi, \Psi)$  be initial data on the singularity and let V be an admissible potential. Let  $(M, g, \varphi)$  solve Einstein's equations and suppose there is a diffeomorphism

 $F:(0, T) \times \Sigma \rightarrow U \subset M$ 

such that

$$F^*g = -dt \otimes dt + h$$
,

 $\operatorname{tr}_h k \to \infty$  as  $t \to 0$ , and

$$\|\mathcal{H}-\mathring{\mathcal{H}}\|_{C^m(\Sigma)}+\|\mathcal{K}-\mathring{\mathcal{K}}\|_{C^m(\Sigma)}+\|\Phi-\mathring{\Phi}\|_{C^m(\Sigma)}+\|\Psi-\mathring{\Psi}\|_{C^m(\Sigma)}\leq C_mt^{\delta}.$$

Then  $(M, g, \varphi)$  is a locally Gaussian development of the data.



Given initial data on the singularity, there exists a locally Gaussian development of the data.



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#### Theorem

Let  $(M, g, \varphi)$  be a locally Gaussian development of the data. Then

$$\|t(\operatorname{tr}_h k) - 1\|_{C^m(\Sigma)} + \|\ln(\operatorname{tr}_h k) + \ln t\|_{C^m(\Sigma)} \leq C_m t^{\delta}.$$

Moreover,

$$F^*g = -dt \otimes dt + \sum_{i,k} b_{ik} t^{2p_{\max\{i,k\}}} \omega^i \otimes \omega^k$$

on (0, T)  $\times \Sigma$ , where  $b_{ik} \in C^{\infty}((0, T) \times \Sigma)$  satisfy  $b_{ik} \rightarrow \delta_{ik}$  as  $t \rightarrow 0$ .



Suppose there are two locally Gaussian developments of the same initial data on the singularity. Then the developments are locally isometric in a neighborhood of the singularity.



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Given initial data on the singularity, there is a maximal locally Gaussian globally hyperbolic development of the data which is unique up to isometry.

## Remark

This is analogous to the classical result by Choquet-Bruhat and Geroch!



- Einstein's equations have a geometric singular initial value problem formulation.
- What about other types of foliations?
- From regular initial data to data on the singularity?



# THANK YOU!