

The Maxwell-scalar field system near spatial infinity

Marica Minucci

Queen Mary University of London

joint work with Dr. Juan Antonio Valiente Kroon

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Outline

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The cylinder at spatial infinity and the F-gauge



The standard representation of spatial infinity

The Minkowski metric $\tilde{\eta}$ written in Cartesian coordinates $(\tilde{x}^\mu) = (\tilde{t}, \tilde{x}^\alpha)$,

$$\tilde{\eta} = \eta_{\mu\nu} d\tilde{x}^\mu \otimes d\tilde{x}^\nu,$$

where $\eta_{\mu\nu} = \text{diag}(1, -1, -1, -1)$. By introducing polar coordinates defined by $\tilde{\rho}^2 \equiv \delta_{\alpha\beta} \tilde{x}^\alpha \tilde{x}^\beta$ where $\delta_{\alpha\beta} = \text{diag}(1, 1, 1)$, and an arbitrary choice of coordinates on \mathbb{S}^2 , the metric $\tilde{\eta}$ can be written as

$$\tilde{\eta} = d\tilde{t} \otimes d\tilde{t} - d\tilde{\rho} \otimes d\tilde{\rho} - \tilde{\rho}^2 \sigma,$$

with $\tilde{t} \in (-\infty, \infty)$, $\tilde{\rho} \in [0, \infty)$ and where σ denotes the standard metric on \mathbb{S}^2 . A conformal representation of the Minkowski spacetime close to i^0 is to make use of *inversion coordinates* $(x^\alpha) = (t, x^i)$ defined by

$$x^\mu = -\tilde{x}^\mu / \tilde{X}^2, \quad \tilde{X}^2 \equiv \tilde{\eta}_{\mu\nu} \tilde{x}^\mu \tilde{x}^\nu,$$

which is valid in the domain

$$\tilde{\mathcal{D}} \equiv \{p \in \mathbb{R}^4 \mid \eta_{\mu\nu} \tilde{x}^\mu(p) \tilde{x}^\nu(p) < 0\}.$$

Using these coordinates the conformal representation of the Minkowski spacetime with *unphysical metric* is given by

$$\bar{\eta} = \Xi^2 \tilde{\eta}, \quad \Xi \equiv X^2.$$

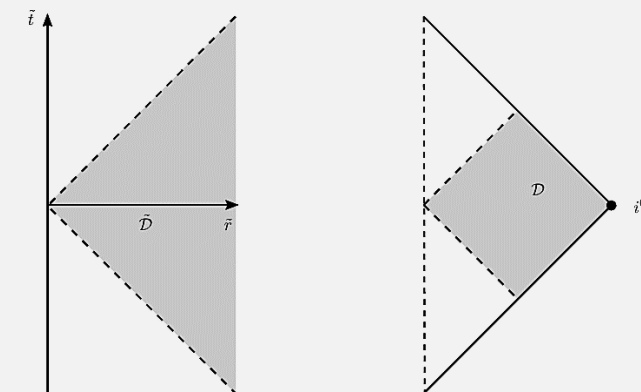
Introducing an *unphysical polar coordinate* via the relation $\rho^2 \equiv \delta_{\alpha\beta} x^\alpha x^\beta$, one finds that the metric $\bar{\eta}$ can be written as

$$\bar{\eta} = dt \otimes dt - d\rho \otimes d\rho - \rho^2 \sigma, \quad \Xi = t^2 - \rho^2,$$

with $t \in (-\infty, \infty)$ and $\rho \in (0, \infty)$. In this conformal representation, spatial infinity i^0 corresponds to the origin of the domain

$$\mathcal{D} \equiv \{p \in \mathbb{R}^4 \mid \eta_{\mu\nu} x^\mu(p) x^\nu(p) < 0\}.$$

This region contains the asymptotic region of the Minkowski spacetime around spatial infinity.



The F-gauge and the conformal structure

The F-gauge

Introducing a time coordinate τ through the relation $t = \rho\tau$ one finds that the metric $\bar{\eta}$ can be written as

$$\bar{\eta} = \rho^2 d\tau \otimes d\tau - (1 - \tau^2) d\rho \otimes d\rho + \rho\tau d\rho \otimes d\tau + \rho\tau d\tau \otimes d\rho - \rho^2 \sigma.$$

The conformal representation containing the *cylinder at spatial infinity* is obtained by considering the rescaled metric

$$\eta \equiv \frac{1}{\rho^2} \bar{\eta}.$$

Introducing the coordinate $\varrho \equiv -\ln \rho$ the metric η can be reexpressed as

$$\eta = d\tau \otimes d\tau - (1 - \tau^2) d\varrho \otimes d\varrho - \tau d\tau \otimes d\varrho - \tau d\varrho \otimes d\tau - \sigma.$$

The conformal structure

Given the conformal extension (\mathcal{M}, η) where

$$\eta = \Theta^2 \tilde{\eta}, \quad \Theta \equiv \rho(1 - \tau^2),$$

and

$$\mathcal{M} \equiv \{p \in \mathbb{R}^4 \mid -1 \leq \tau \leq 1, \rho(p) \geq 0\}.$$

In this representation future and past null infinity are described by the sets

$$\mathcal{I}^+ \equiv \{p \in \mathcal{M} \mid \tau(p) = 1\}, \quad \mathcal{I}^- \equiv \{p \in \mathcal{M} \mid \tau(p) = -1\},$$

while the physical Minkowski spacetime can be identified with the set

$$\tilde{\mathcal{M}} \equiv \{p \in \mathcal{M} \mid -1 < \tau(p) < 1, \rho(p) > 0\},$$

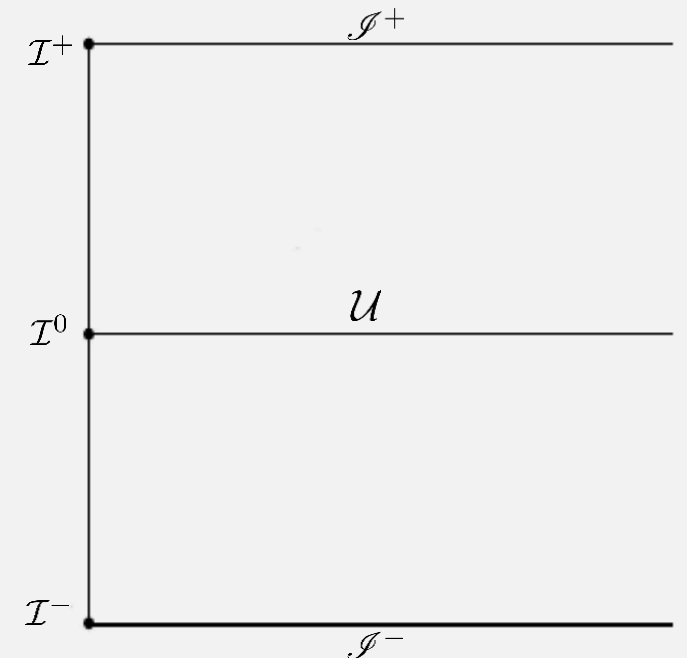
In addition, the *cylinder at spatial infinity* is given by

$$\mathcal{I} \equiv \{p \in \mathcal{M} \mid |\tau(p)| < 1, \rho(p) = 0\},$$

and *critical sets* are

$$\mathcal{I}^+ \equiv \{p \in \mathcal{M} \mid \tau(p) = 1, \rho(p) = 0\}, \quad \mathcal{I}^- \equiv \{p \in \mathcal{M} \mid \tau(p) = -1, \rho(p) = 0\}.$$

Spatial infinity i^0 , which is at infinity respect to the metric η , corresponds to a set which has the topology of $\mathbb{R} \times \mathbb{S}^2$.



The Maxwell-scalar field system

Conformal properties

Let \tilde{F}_{ab} denote the *Faraday tensor* over a spacetime $(\tilde{\mathcal{M}}, \tilde{g})$ and let $\tilde{\nabla}$ be the Levi-Civita connection of the metric \tilde{g} . The Maxwell equations with source are given by

$$\begin{aligned}\tilde{\nabla}^a \tilde{F}_{ab} &= \tilde{J}_b, \\ \tilde{\nabla}_{[a} \tilde{F}_{bc]} &= 0.\end{aligned}$$

The homogeneous equation is automatically satisfied if one sets

$$\tilde{F}_{ab} = \tilde{\nabla}_a \tilde{A}_b - \tilde{\nabla}_b \tilde{A}_a.$$

Coupled to the above, we consider the *conformally invariant wave equation*

$$\tilde{\mathfrak{D}}_a \tilde{\mathfrak{D}}^a \tilde{\phi} - \frac{1}{6} \tilde{R} \tilde{\phi} = 0,$$

The coupling between the Maxwell field and the scalar field is encoded in the covariant derivative

$$\tilde{\mathfrak{D}}_a = \tilde{\nabla}_a - iq \tilde{A}_a.$$

The current \tilde{J}_a is given by

$$\tilde{J}_a = iq \left(\tilde{\phi} \tilde{\mathfrak{D}}_a \tilde{\phi} - \tilde{\phi} \overline{(\tilde{\mathfrak{D}}_a \tilde{\phi})} \right)$$

Consider a conformal rescaling of the form

$$g_{ab} = \Xi^2 \tilde{g}_{ab}.$$

Associated to the latter we define the *unphysical Faraday tensor*, *unphysical vector potential* and the *unphysical scalar field* via

$$F_{ab} \equiv \tilde{F}_{ab}, \quad A_a \equiv \tilde{A}_a, \quad \phi \equiv \Xi^{-1} \tilde{\phi},$$

Let $F_{AA'BB'}$ denote the spinorial counterpart of the Faraday tensor F_{ab} . It satisfies the well-know decomposition

$$F_{AA'BB'} = \phi_{AB}\epsilon_{A'B'} + \bar{\phi}_{A'B'}\epsilon_{AB},$$

where $\phi_{AB} = \phi_{(AB)}$ is the so-called *Maxwell spinor*. The unphysical Maxwell equations are equivalent to

$$\nabla^B{}_{A'}\phi_{AB} = J_{AA'},$$

where

$$J_{AA'} \equiv iq\left(\bar{\phi}\nabla_{AA'}\phi - \phi\nabla_{AA'}\bar{\phi}\right) + 2q^2|\phi|^2A_{AA'}.$$

The study of the Maxwell-scalar field system can be reduced, making use of the generalised Lorenz gauge condition

$$\nabla^{AA'}A_{AA'} = 0,$$

to the system of wave equations

$$\text{(A)} \quad \square\phi = q^2\phi A_{AA'}A^{AA'} + 2iqA^{AA'}\nabla_{AA'}\phi,$$

$$\text{(B)} \quad \square A_{AA'} + 2\Phi_{ABA'B'}A^{BB'} = 2q|\phi|^2A_{AA'} + iq\bar{\phi}\nabla_{AA'}\phi - iq\phi\nabla_{AA'}\bar{\phi}.$$

These equations are supplemented by initial conditions for the values of ϕ and $A_{AA'}$ and of their normal derivatives.

Decomposition of the equations in the space-spinor formalism and the Newman-Penrose frame



Space-spinor decomposition

Let $\tau^{AA'}$ denote the spinorial counterpart of a timelike vector field τ^a tangent to a congruence of curves. The Hermitian spinor $\tau^{AA'}$ is chosen to have the normalisation

$$\tau_{AA'}\tau^{AA'} = 2.$$

Consistent with the latter, we consider a spin dyad $\{o^A, \iota^A\}$ chosen so that

$$\tau^{AA'} = o^A\bar{o}^{A'} + \iota^A\bar{\iota}^{A'}.$$

It follows then that

$$\tau_{AA'}\tau^{BA'} = \delta_A^B.$$

The above relations induce a Hermitian conjugation operation via the relation

$$\mu_A^\dagger \equiv \tau_A^{A'}\bar{\mu}_{A'}.$$

The Newman-Penrose frame

The Newman-Penrose (NP) frame satisfies

$$g(e_{AA'}, e_{BB'}) = \epsilon_{AB}\epsilon_{A'B'},$$

of the form

$$\begin{aligned} e_{00'} &= \frac{1}{\sqrt{2}} \left((1 - \tau)\partial_\tau + \rho\partial_\rho \right), \\ e_{11'} &= \frac{1}{\sqrt{2}} \left((1 + \tau)\partial_\tau - \rho\partial_\rho \right), \\ e_{01'} &= -\frac{1}{\sqrt{2}} \mathbf{X}_+, \\ e_{10'} &= -\frac{1}{\sqrt{2}} \mathbf{X}_-, \end{aligned}$$

where \mathbf{X}_+ and \mathbf{X}_- are vectors spanning the tangent space of S^2 with dual covectors α_+ and α_- such that metric of the 2-sphere is given by

$$\sigma = 2(\alpha^+ \otimes \alpha^- + \alpha^- \otimes \alpha^+).$$

We define the *F-reduced wave operator* \blacksquare acting on a scalar ζ as

$$\blacksquare\zeta \equiv (1 - \tau^2)\ddot{\zeta} + 2\tau\rho\dot{\zeta}' - \rho^2\zeta'' - 2\tau\dot{\zeta} - \frac{1}{2}\bar{\partial}\bar{\partial}\zeta - \frac{1}{2}\bar{\partial}\bar{\partial}\zeta,$$

After some computations carried out using the suite `xAct` for tensorial and spinorial manipulations in the `Wolfram` programming language, the wave equations **(A)** and **(B)** are equivalent to the scalar system:

$$\begin{aligned} \text{(A')} \quad & \blacksquare\phi = s, \\ \text{(B')} \quad & \blacksquare\alpha_+ - 2\dot{\alpha}_+ = j_+, \\ \text{(C')} \quad & \blacksquare\alpha_- + 2\dot{\alpha}_- = j_-, \\ \text{(D')} \quad & \blacksquare\alpha_0 + \alpha_0 = j_0, \\ \text{(E')} \quad & \blacksquare\alpha_2 + \alpha_2 = j_2. \end{aligned}$$

The scalars α , α_0 , α_1 and α_2 have spin weight 0, -1 , 0, 1 whereas α_{\pm} and j_{\pm} are defined by

$$\alpha_{\pm} \equiv \alpha \pm \alpha_1, \quad j_{\pm} \equiv j \pm j_1,$$

all of them of spin-weight 0.

Key observation: the cylinder at infinity \mathcal{I} is a *total characteristic* of the Maxwell-scalar field system.

This property can be used to construct, in a recursive manner, the *jets of order p at \mathcal{I}* , $J^p[\phi, \alpha]$, $p \geq 0$ of the solutions to the evolution equations.

The jet is defined as

$$J^p[\phi, \alpha] \equiv \{\partial_{\rho}^p \phi|_{\rho=0}, \partial_{\rho}^p \alpha|_{\rho=0}\}.$$

The F-reduced wave operator \blacksquare , reduces to an operator intrinsic to \mathcal{I} . Defining the *F-reduced wave operator on \mathcal{I}* , $\blacktriangle \equiv \blacksquare|_{\mathcal{I}}$, acting on a scalar ζ as

$$\blacktriangle\zeta \equiv (1 - \tau^2)\ddot{\zeta} - 2\tau\dot{\zeta} - \frac{1}{2}(\bar{\partial}\bar{\partial} + \bar{\partial}\bar{\partial})\zeta.$$

The decoupled case

Setting the charge parameter $\mathfrak{q} = 0$, the equations reduce to the linear system of equations

- $\phi = 0$
- $\alpha_+ = 0,$
- $\alpha_- = 0,$
- $\alpha_0 + 2\dot{\alpha}_0 = 0,$
- $\alpha_2 - 2\dot{\alpha}_2 = 0.$

Defining

$$\phi^{(p)} \equiv \partial_\rho^p \phi|_{\rho=0}, \quad \alpha_\pm^{(p)} \equiv \partial_\rho^p \alpha_\pm|_{\rho=0}, \quad \alpha_0^{(p)} \equiv \partial_\rho^p \alpha_0|_{\rho=0}, \quad \alpha_2^{(p)} \equiv \partial_\rho^p \alpha_2|_{\rho=0}, \quad p \geq 0,$$

the order elements of $J^p[\phi, \boldsymbol{\alpha}]$ solutions satisfy a set of equations that can be summarised as three model equations:

- (a) $\blacktriangle \zeta + 2p\tau \dot{\zeta} = 0,$
- (b) $\blacktriangle \zeta + 2(p\tau + 1)\dot{\zeta} = 0,$
- (c) $\blacktriangle \zeta + 2(p\tau - 1)\dot{\zeta} = 0.$

It is assumed that:

The scalar ζ in equation **(a)** has spin-weight 0 and admits an expansion of form

$$\zeta = \sum_{p=0}^{\infty} \sum_{l=0}^p \sum_{m=-1}^l \frac{1}{p!} \zeta_{p;l,m} (Y_{lm}) \rho^p;$$

The scalar ζ in equation **(b)** has spin-weight 1 and admits an expansion of the form

$$\zeta = \sum_{p=1}^{\infty} \sum_{l=1}^p \sum_{m=-l}^l \frac{1}{p!} \zeta_{0p;l,m} ({}_1Y_{lm}) \rho^p;$$

The scalar ζ in equation **(c)** has spin-weight -1 and an expansion of the form

$$\zeta = \sum_{p=1}^{\infty} \sum_{l=1}^p \sum_{m=-l}^l \frac{1}{p!} \zeta_{p;l,m} ({}_{-1}Y_{lm}) \rho^p.$$

Substituting the Ansatz into the model equations **(a)**-**(c)** one obtains the ordinary differential equations

$$\begin{aligned} \text{(a)} \quad & (1 - \tau^2) \ddot{\zeta}_{p;\ell,m} + 2\tau(p-1) \dot{\zeta}_{p;\ell,m} + (p+\ell)(\ell-p+1) \zeta_{p;\ell,m} = 0, \\ \text{(b)} \quad & (1 - \tau^2) \ddot{\zeta}_{0,p;\ell,m} + 2((p-1)\tau + 1) \dot{\zeta}_{0,p;\ell,m} + (p+\ell)(\ell-p+1) \zeta_{0,p;\ell,m} = 0, \\ \text{(c)} \quad & (1 - \tau^2) \ddot{\zeta}_{2,p;\ell,m} + 2((p-1)\tau - 1) \dot{\zeta}_{2,p;\ell,m} + (p+\ell)(\ell-p+1) \zeta_{2,p;\ell,m} = 0. \end{aligned}$$

It can be verified that if $\zeta_{0,p;p,m}(\tau)$ is a solution to equation **(b)** then $\zeta^s_{0,p;p,m}(\tau) \equiv \zeta_{0,p;p,m}(-\tau)$ solves **(c)**. Thus, it is only necessary to study two model equations.

These equations are examples of *Jacobi ordinary differential equations*.

The behaviour of the solutions to equations **(a)**-**(c)** depends on the value of the parameter ℓ

For $0 \leq \ell \leq p - 1$, $p \geq 1$ the nature of the solutions is summarised in the following:

lemma

The solutions to the system **(a)**-**(c)** can be written as

$$\zeta_{p;\ell,m}(\tau) = A_{p;\ell,m} \left(\frac{1-\tau}{2} \right)^p P_\ell^{(p,-p)}(\tau) + B_{p;\ell,m} \left(\frac{1+\tau}{2} \right)^p P_\ell^{(-p,p)}(\tau),$$

$$\zeta_{0,p;\ell,m}(\tau) = C_{p;\ell,m} \left(\frac{1-\tau}{2} \right)^{(p+1)} P_\ell^{(1+p,1-p)}(\tau) + D_{p;\ell,m} \left(\frac{1+\tau}{2} \right)^{(p-1)} P_\ell^{(-1-p,p-1)}(\tau),$$

$$\zeta_{2,p;\ell,m}(\tau) = E_{p;\ell,m} \left(\frac{1-\tau}{2} \right)^{(p+1)} P_\ell^{(1+p,1-p)}(-\tau) + F_{p;\ell,m} \left(\frac{1+\tau}{2} \right)^{(p-1)} P_\ell^{(-1-p,p-1)}(-\tau),$$

with $P_n^{(\alpha,\beta)}(\tau)$ Jacobi polynomials of order n and where

$$A_{p;\ell,m}, \quad B_{p;\ell,m}, \quad C_{p;\ell,m}, \quad D_{p;\ell,m}, \quad E_{p;\ell,m}, \quad F_{p;\ell,m} \in \mathbb{C}$$

denote some constants which can be expressed in terms of the initial conditions.

However, for equation **(a)** in the case $\ell = p$ we have the following lemma:

lemma

For $l = p$ the solution to the equation **(a)** can be written as

$$\zeta_{p;p,m}(\tau) = \left(\frac{1-\tau}{2} \right)^p \left(\frac{1+\tau}{2} \right)^p \left(C_{1,p;\ell,m} + C_{2,p;\ell,m} \int_0^\tau \frac{ds}{(1-s^2)^{p+1}} \right),$$

where $C_{1,p;\ell,m}$, $C_{2,p;\ell,m}$ are integration constants.

Letting $\zeta_{\star p;p,m} \equiv \zeta_{p;p,m}(0)$ and $\dot{\zeta}_{\star p;p,m} \equiv \dot{\zeta}_{p;p,m}(0)$ one finds that there is no logarithmic divergence if and only if $\dot{\zeta}_{\star p;p,m} = 0$ —that is, *when the initial data for ζ is time symmetric*.

Similarly, equations (b) and (c) one obtains an analogous result:

lemma

For $\ell = p$ the solution to equations (b) and (c) can be written as

$$\zeta_{0;p;p,m}(\tau) = \left(\frac{1-\tau}{2}\right)^{(p+1)} \left(\frac{1+\tau}{2}\right)^{(p-1)} \left(C_{3;p;\ell,m} + C_{4;p;\ell,m} \int_0^\tau \frac{ds}{(1-s)^{p+2}(1+s)^p} \right),$$

$$\zeta_{2;p;p,m}(\tau) = \left(\frac{1-\tau}{2}\right)^{(p-1)} \left(\frac{1+\tau}{2}\right)^{(p+1)} \left(C_{5;p;\ell,m} + C_{6;p;\ell,m} \int_0^\tau \frac{ds}{(1+s)^{p+2}(1-s)^p} \right),$$

where $C_{3;p;\ell,m}$, $C_{4;p;\ell,m}$, $C_{5;p;\ell,m}$, $C_{6;p;\ell,m}$ are integration constants.

If we let $\zeta_{0,\star;p;p,m} \equiv \zeta_{0;p;p,m}(0)$, $\dot{\zeta}_{0,\star;p;p,m} \equiv \dot{\zeta}_{0;p;p,m}(0)$ and $\dot{\zeta}_{0\star;p;p,m} = -2\zeta_{0\star;p;p,m}$ there is no logarithmic divergence.

Now, since $a_s(\tau) \equiv a(-\tau)$ is a solution for the equation for $\zeta_{2;p;p,m}$ we have that to avoid logarithms in the solutions to equation (c) one needs the condition

$$\dot{\zeta}_{2\star;p;p,m} = 2\zeta_{2\star;p;p,m}.$$

Proposition

Given the jet $J^p[\phi, \alpha]$ for $\mathfrak{q} = 0$ one has that:

- (i) the elements of the jet have polynomial dependence in τ for the harmonic sectors with $0 \leq \ell \leq p - 1$ and, thus, they extend analytically through $\tau = \pm 1$;
- (ii) generically, for $\ell = p$, the solutions have logarithmic singularities at $\tau = \pm 1$. These logarithmic divergences can be precluded by fine-tuning of the initial data.

The coupled case

The analysis of the case $\mathfrak{q} \neq 0$ begins by looking at the solutions corresponding to the jet $J^0[\phi, \alpha]$ – that is $p = 0$.

The $p = 0$ order transport equations

Evaluating the system one finds that

$$\begin{aligned} \text{(a')} \quad & \blacktriangle \phi^{(0)} = s^{(0)}, \\ \text{(b')} \quad & \blacktriangle \alpha_+^{(0)} - 2\dot{\alpha}_+^{(0)} = j_+^{(0)}, \\ \text{(c')} \quad & \blacktriangle \alpha_-^{(0)} + 2\dot{\alpha}_-^{(0)} = j_-^{(0)}, \\ \text{(d')} \quad & \blacktriangle \alpha_0^{(0)} + \alpha_0^{(0)} = j_0^{(0)}, \\ \text{(e')} \quad & \blacktriangle \alpha_2^{(0)} + \alpha_2^{(0)} = j_2^{(0)}. \end{aligned}$$

These transport equations decouple and it is possible to write down the solution explicitly. More precisely, one has that:

lemma

The unique solution to the 0-th order system (a')-(e') with suitable initial conditions is given by

$$\phi^{(0)} = \varphi_*, \quad \alpha_{\pm}^{(0)} = 0, \quad \alpha_0^{(0)} = 0, \quad \alpha_2^{(0)} = 0.$$

The $p \geq 1$ transport equations

In order to analyse the properties of the jet of order p , $J^p[\phi, \boldsymbol{\alpha}]$ for given $p = n$, we assume that we have knowledge of the jets

$$J^0[\phi, \boldsymbol{\alpha}], J^1[\phi, \boldsymbol{\alpha}], \dots, J^{n-1}[\phi, \boldsymbol{\alpha}].$$

Under this assumption and taking into account the previous Lemma it follows that the elements of $J^p[\phi, \boldsymbol{\alpha}]$ satisfy the equations:

$$\begin{aligned} \text{(a'')} \quad & \blacktriangle \phi^{(n)} + 2n\tau \dot{\phi}^{(n)} = s^{(n)}, \\ \text{(b'')} \quad & \blacktriangle \alpha_+^{(n)} + 2(n\tau - 1)\dot{\alpha}_+^{(n)} = 2q^2|\varphi_\star|^2 \alpha_+^{(n)} + \tilde{j}_+^{(n)}, \\ \text{(c'')} \quad & \blacktriangle \alpha_-^{(n)} + 2(n\tau + 1)\dot{\alpha}_-^{(n)} = 2q^2|\varphi_\star|^2 \alpha_-^{(n)} + \tilde{j}_-^{(n)}, \\ \text{(d'')} \quad & \blacktriangle \alpha_0^{(n)} + 2n\tau \dot{\alpha}_0^{(n)} + \alpha_0^{(n)} = 2q^2|\varphi_\star|^2 \alpha_0^{(n)} + \tilde{j}_0^{(n)}, \\ \text{(e'')} \quad & \blacktriangle \alpha_2^{(n)} + 2n\tau \dot{\alpha}_2^{(n)} + \alpha_2^{(n)} = 2q^2|\varphi_\star|^2 \alpha_2^{(n)} + \tilde{j}_2^{(n)}, \end{aligned}$$

where $s^{(n)}$, $\tilde{j}_\pm^{(n)}$, $\tilde{j}_0^{(n)}$ and $\tilde{j}_2^{(n)}$ depend, solely, on the elements of $J^p[\phi, \boldsymbol{\alpha}]$, $0 \leq p \leq n - 1$.

In particular, it is possible to consider model homogeneous equations of the form

$$\blacktriangle \zeta + 2(n\tau - 1)\dot{\zeta} - \varkappa \zeta = 0, \tag{1a}$$

$$\blacktriangle \zeta + 2n\tau \dot{\zeta} + (1 - \varkappa)\zeta = 0 \tag{1b}$$

with $\varkappa \equiv 2q^2|\varphi_\star|^2$. The solutions of these equations for generic choice of \varkappa is radically different to that of the case $\varkappa = 0$ —i.e. $q = 0$.

Assuming that the various fields have an asymptotic expansion as before, one is led to consider a hierarchy of ordinary differential equations of the form

$$\begin{aligned}
 \text{(a'')} \quad & (1 - \tau^2)\ddot{\phi}_{n;\ell,m} + 2(n-1)\tau\dot{\phi}_{n;\ell,m} + ((\ell - n + 1)(n + \ell))\phi_{n;\ell,m} = s_{n;\ell,m}, \\
 \text{(b'')} \quad & (1 - \tau^2)\ddot{\alpha}_{+,n;\ell,m} + 2(-1 + (n-1)\tau)\dot{\alpha}_{+,n;\ell,m} + (\ell(\ell + 1) - n(n-1) - \varkappa)\alpha_{+,n;\ell,m} = \tilde{j}_{+,n;\ell,m}, \\
 \text{(c'')} \quad & (1 - \tau^2)\ddot{\alpha}_{-,n;\ell,m} + 2(1 + (n-1)\tau)\dot{\alpha}_{-,n;\ell,m} + (\ell(\ell + 1) - n(n-1) - \varkappa)\alpha_{-,n;\ell,m} = \tilde{j}_{-,n;\ell,m}, \\
 \text{(d'')} \quad & (1 - \tau^2)\ddot{\alpha}_{0,n;\ell,m} + 2(n-1)\tau\dot{\alpha}_{0,n;\ell,m} + ((\ell - n + 1)(n + \ell) - \varkappa)\alpha_{0,n;\ell,m} = \tilde{j}_{0,n;\ell,m}, \\
 \text{(e'')} \quad & (1 - \tau^2)\ddot{\alpha}_{2,n;\ell,m} + 2(n-1)\tau\dot{\alpha}_{2,n;\ell,m} + ((\ell - n + 1)(n + \ell) - \varkappa)\alpha_{2,n;\ell,m} = \tilde{j}_{2,n;\ell,m},
 \end{aligned}$$

for $0 \leq \ell \leq n$, $-\ell \leq m \leq \ell$ and with the *source terms*

$$s_{n;\ell,m}, \quad \tilde{j}_{+,n;\ell,m}, \quad \tilde{j}_{-,n;\ell,m}, \quad \tilde{j}_{0,n;\ell,m}, \quad \tilde{j}_{2,n;\ell,m},$$

known as a result of the spherical harmonics decomposition of the lower order jets $J^p[\phi, \alpha]$ for $0 \leq p \leq n - 1$.

For this, Frobenius's method is resorted to study the properties of the equations in terms of asymptotic expansions at the values $\tau = \pm 1$. The homogeneous version of equation (a'')-(e'') can be described in terms of the model equation

$$(1 - \tau^2)\ddot{\zeta} + 2(\varsigma + (n-1)\tau)\dot{\zeta} + (\ell(\ell + 1) - n(n-1) - \varkappa)\zeta = 0$$

where

$$\varsigma = \begin{cases} -1 & \text{for } \alpha_+ \\ 1 & \text{for } \alpha_- \\ 0 & \text{for } \phi, \alpha_0, \alpha_2 \end{cases}$$

—recall also that $\varkappa = 2q^2\varphi_*^2$.

Following Frobenius's method we look for power series solutions of the form

$$\zeta = (1 - \tau)^r \sum_{k=0}^{\infty} D_k (1 - \tau)^k, \quad D_0 \neq 0.$$

Substitution of this Ansatz into the model equation leads to the *indicial equation*

$$2r(r - 1) - 2\varsigma r - 2(n - 1)r = 0.$$

ς	r_1	r_2
-1	0	$n - 1$
0	0	n
1	0	$n + 1$

Once the solutions to the indicial equation are known, the Ansatz leads to a recurrence relation for the coefficients D_k in the series.

The root $r_1 = 0$ of the indicial polynomial does not lead to a valid series solution.

We look for a second solution of the form

$$\zeta = \sum_{k=0}^{\infty} G_k (1 - \tau)^k + (1 - \tau)^{r_2} \log(1 - \tau) \sum_{k=0}^{\infty} M_k (1 - \tau)^k, \quad G_0 \neq 0, \quad M_0 \neq 0.$$

The recurrence relations implied by the Ansatz shows that all the coefficients M_k for $k = 1, 2, \dots$ and G_k for $k = 0, 1, 2, \dots$ can be expressed in terms of the coefficient M_0 .

Proposition

The general solution to

$$(1 - \tau^2)\ddot{\zeta} + 2(\varsigma + (n - 1)\tau)\dot{\zeta} + (\ell(\ell + 1) - n(n - 1) - \varkappa)\zeta = 0, \quad \varkappa \neq 0$$

with $\varsigma = -1, 0, 1$ and $0 \leq \ell \leq n$, $n = 1, 2, \dots$ consists, of:

(i) one solution which is analytic for $\tau \in [-1, 1]$;

(ii) one solution which is analytic for $\tau \in (-1, 1)$ and has logarithmic singularities at $\tau = \pm 1$. At these singular points the solution is of class C^{r_2-1} .

Main theorem

For generic data for Maxwell-scalar field system which have finite energy and are analytic around \mathcal{I} , the solution to the transport equations on \mathcal{I} develop logarithmic singularities at the critical points \mathcal{I}^+ and \mathcal{I}^- .

The nonlinear interaction makes both fields more singular than what is seen when the fields are non-interacting!

Thank you for your attention